Series representation of a function. The main purpose of series is to write a given complicated quantity as an infinite sum of simple terms; and since the terms get smaller and smaller, we can approximate the original quantity by taking only the first few terms of the series. In this section, we finally develop the tool that lets us do this in most cases: a way to write any reasonable function as an explicit power series. This will allow us to compute outputs of the function by plugging into the series.

Our functions must behave decently near the center point of the desired power series. We say f(x) is analytic at x=a if it is possible to write $f(x)=\sum_{n=0}^{\infty}c_n(x-a)^n$ for some coefficients c_n , with positive radius of convergence. In practice, any formula involving standard functions and operations defines an analytic function, provided the formula gives real number values in a small interval around x=a. For example $\frac{1}{x-a}$ is not analytic at x=a, because it gives $\pm \infty$ at x=a; and $\sqrt{x-a}$ is not analytic at x=a because for x slightly smaller than a, it gives the square root of a negative number.*

Taylor Series Theorem: Let f(x) be a function which is analytic at x = a. Then we can write f(x) as the following power series, called the Taylor series of f(x) at x = a:

$$f(x) = f(a) + f'(a)(x-a) + \frac{f''(a)}{2!}(x-a)^2 + \frac{f'''(a)}{3!}(x-a)^3 + \frac{f''''(a)}{4!}(x-a)^4 + \cdots,$$

valid for x within a radius of convergence |x-a| < R with R > 0, or convergent for all x.

If we write the nth derivative of f(x) as $f^{(n)}(x)$, this becomes:

$$f(x) = \sum_{n=0}^{\infty} c_n (x-a)^n$$
 with coefficients $c_n = \frac{f^{(n)}(a)}{n!}$.

WARNING: The coefficients are *constants* with no x, so $c_1 = f'(a)$, NOT f'(x).

Proof. By hypothesis f(x) is analytic, so $f(x) = \sum_{n=0}^{\infty} c_n(x-a)^n$ for some c_n ; we will derive the desired formula for these coefficients. Since $f(a) = \sum_{n=0}^{\infty} c_n(a-a)^n = c_0 + c_1(0) + c_2(0^2) + \cdots$, we get $c_0 = f(a)$. Next, by the Theorem in §11.9, we have $f'(x) = \sum_{n=0}^{\infty} nc_n(x-a)^{n-1}$, so $f'(a) = c_1 + 2c_2(0) + 3c_3(0^2) + \cdots$, and $c_1 = f'(a)$. Next, $f''(x) = \sum_{n=0}^{\infty} n(n-1)c_n(x-a)^{n-2}$, so $f''(a) = (2)(1)c_2$ and $c_2 = \frac{1}{2}f''(a)$. Continuing, we get:

$$f^{(N)}(x) = \sum_{n=1}^{\infty} n(n-1) \cdots (n-N+1) c_n (x-a)^{n-N}$$
.

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^{*}The function $\sqrt[3]{x-a}$ is also not analytic near x=a, ever though it gives real number values. The problem is that it has a vertical tangent at x=a, so it is not differentiable.

The terms for n = 0, 1, ..., N-1 are all zero because of the factors $n(n-1) \cdots (n-N+1)$, so the first non-zero term is for n = N. Plugging in x = a gives: $f^{(N)}(a) = N(N-1) \cdots (1)c_N$, and $c_N = \frac{1}{N!} f^{(N)}(a)$ as desired. Q.E.D.

Once we have a power series for f(x) with known coefficients $c_n = \frac{f^{(n)}(a)}{n!}$, we can approximate f(x) by taking a finite partial sum of the series up to some cutoff term N. This partial sum is called a Taylor polynomial, denoted $T_N(x)$:

$$f(x) \approx T_N(x) = \sum_{n=0}^N c_n(x-a)^n = f(a) + f'(a)(x-a) + \dots + \frac{f^{(N)}(a)}{N!}(x-a)^N.$$

Note that $T_1(x) = f(a) + f'(a)(x - a)$ is just the linear approximation near x = a, whose graph is the tangent line (Calculus I §2.9). We can improve this approximation of f(x) in two ways:

- Take more terms, increasing N.
- Take the center a close to x, giving small (x-a) and tiny $(x-a)^n$.

A Taylor series centered at a = 0 is specially named a Maclaurin series.

Example: sine function. To find Taylor series for a function f(x), we must determine $f^{(n)}(a)$. This is easiest for a function which satisfies a simple differential equation relating the derivatives to the original function. For example, $f(x) = \sin(x)$ satisfies f''(x) = -f(x), so coefficients of the Maclaurin series (center a = 0) are:

n	0	1	2	3	4	5	6	7
$f^{(n)}(x)$	$\sin(x)$	$\cos(x)$	$-\sin(x)$	$-\cos(x)$	$\sin(x)$	$\cos(x)$	$-\sin(x)$	$-\cos(x)$
$f^{(n)}(0)$	0	1	0	-1	0	1	0	-1
$c_n = \frac{f^{(n)}(0)}{n!}$	0	1	0	$-\frac{1}{3!}$	0	1 5!	0	$-\frac{1}{7!}$

That is:

$$\sin(x) = x - \frac{x^3}{3!} + \frac{x^5}{5!} - \frac{x^7}{7!} + \dots = \sum_{n=0}^{\infty} (-1)^n \frac{x^{2n+1}}{(2n+1)!}.$$

To find the domain of convergence, we apply the Ratio Test (11.6/I):

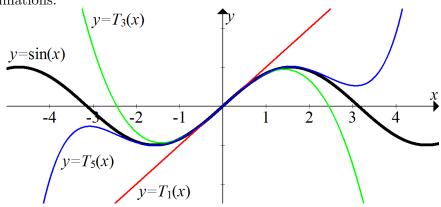
$$L = \lim_{n \to \infty} \left| \frac{a_{n+1}}{a_n} \right| = \lim_{n \to \infty} \left| \frac{x^{2(n+1)+1}}{(2(n+1)+1)!} \right| / \frac{x^{2n+1}}{(2n+1)!}$$

$$= \lim_{n \to \infty} \frac{|x|^{2n+3}}{|x|^{2n+1}} \cdot \frac{(2n+1)!}{(2n+3)!} = \lim_{n \to \infty} \frac{|x|^2}{(2n+2)(2n+3)} = 0$$

for any fixed $x \neq 0$. Since L = 0 < 1 regardless of x, the series converges for all x.

This formula for sin(x) astonishes because the right side is a simple algebraic series having no apparent relation to trigonometry. We can try to understand and

check the series by graphically comparing sin(x) with its first few Taylor polynomial approximations:



- The Taylor polynomial $T_1(x) = x$ (in red) is just the linear approximation or tangent line of $y = \sin(x)$ at the center point x = 0. The curve and line are close (to within a couple of decimal places) near the point of tangency and up to about $|x| \leq 0.5$. Once they veer apart, the approximation is useless.
- The next Taylor polynomial $T_3(x) = x \frac{x^3}{3!} = x \frac{1}{6}x^3$ (in green) matches $y = \sin(x)$ in its first three derivatives at x = 0, and stays close to the original curve up to about $|x| \le 1.5$.
- The next $T_5(x) = x \frac{x^3}{3!} + \frac{x^5}{5!} = x \frac{1}{6}x^3 + \frac{1}{120}x^5$ is even closer to f(x) for even larger x. Taking enough terms in the Taylor series will give a good approximation for $any \ x$, since the series converges everywhere.

PROBLEM: Compute $\sin(10^{\circ})$. A geometric method would be to construct a right triangle with a 10° angle, and measure the opposite side divided by the hypotenuse; but this would only work for a couple of decimal places of accuracy. Of course, a calculator can produce many decimal places, but how does it know? By Taylor series!

As always when doing calculus on trig functions, we must first convert to radians (see end of §2.5): $10^{\circ} = \frac{2\pi}{360}(10) = \frac{\pi}{18}$. Here $|x| = \frac{\pi}{18} \approx \frac{1}{6}$ is small, so the Maclaurin series centered at 0 should converge quickly, giving very accurate approximations:

$$\sin(\frac{\pi}{18}) \approx T_3(\frac{\pi}{18}) = \frac{\pi}{18} - \frac{1}{6}(\frac{\pi}{18})^3 \approx 0.1736468.$$

It turns out this is correct to 5 decimal places (underlined), using only two non-zero terms of the Taylor series and a good estimate for π . We could verify this by taking more terms and seeing that these 5 digits do not change; or by the Remainder Estimates below.

Example: square roots. We compute $\sqrt{2}$ to 5 decimal places.[†] First, we must consider $\sqrt{2}$ to be an output of the function $f(x) = \sqrt{x}$ at x = 2. Next, we must choose the center a for its Taylor series.

- a=0 does not work because \sqrt{x} is not analytic at x=0. Indeed, if there were a convergent Taylor series $\sqrt{x}=c_0+c_1x+c_2x^2+\cdots$, we could plug in x=-0.1 to get: $\sqrt{-0.1}=c_0+c_1(-0.1)+c_2(-0.1)^2+\cdots$, a real value for the square root of a negative number!
- a = 1 is too far from x = 2: it turns out |x-a| = |2-1| = 1 is beyond the radius of convergence of the Taylor series.
- a=2 is useless, since writing the Taylor series requires us to know $f^{(n)}(2)$, including $f(2)=\sqrt{2}$, the same number we are trying to compute.
- A useful choice of a requires: a > 0 so that the Taylor series exists; a is close to x = 2, making |x-a| small so the series converges quickly; and $f(a) = \sqrt{a}$ is easy to compute so we can find the coefficients. A value satisfying all three conditions is: $a = \frac{9}{4}$.

Now we have:

n	0	1	2	3	4
$f^{(n)}(x)$	$x^{1/2}$	$\frac{1}{2} x^{-1/2}$	$-\frac{1\cdot 1}{2\cdot 2} x^{-3/2}$	$\frac{1\cdot 1\cdot 3}{2\cdot 2\cdot 2} x^{-5/2}$	$-\frac{1\cdot 1\cdot 3\cdot 5}{2\cdot 2\cdot 2\cdot 2} x^{-7/2}$
$f^{(n)}(\frac{9}{4})$	$\frac{3}{2}$	$\frac{1}{3}$	$-\frac{2}{27}$	$\frac{4}{81}$	$-\frac{40}{729}$
$c_n = \frac{f^{(n)}(\frac{9}{4})}{n!}$	$\frac{3}{2}$	$\frac{1}{3}$	$-\frac{1}{27}$	$\frac{2}{243}$	$-\frac{5}{2187}$

Hence:

$$\sqrt{x} = \frac{3}{2} + \frac{1}{3}(x - \frac{9}{4}) - \frac{1}{27}(x - \frac{9}{4})^2 + \frac{2}{243}(x - \frac{9}{4})^3 - \frac{5}{2187}(x - \frac{9}{4})^4 + \cdots$$

$$= \frac{3}{2} + \frac{1}{3}(x - \frac{9}{4}) + \sum_{n=2}^{\infty} (-1)^{n-1} \frac{(2n-3)!!}{n!} \frac{2^{n-1}}{3^{2n-1}} (x - \frac{9}{4})^n,$$

where we use the odd factorial notation $(2n-3)!! = (1)(3)(5)\cdots(2n-3)$. For x=2, we have $x-\frac{9}{4}=-\frac{1}{4}$, so:

$$\sqrt{2} = \frac{3}{2} + \frac{1}{3}(-\frac{1}{4}) - \frac{1}{27}(-\frac{1}{4})^2 + \frac{2}{243}(-\frac{1}{4})^3 - \frac{5}{2187}(-\frac{1}{4})^4 + \cdots,
\approx \frac{3}{2} - \frac{1}{3}\frac{1}{4} - \frac{1}{27}\frac{1}{4^2} - \frac{2}{243}\frac{1}{4^3} - \frac{5}{2187}\frac{1}{4^4} \approx \underline{1.41421}43,$$

which is correct to 5 decimal places (underlined).

[†]We saw another very good algorithm for this in Calculus I §3.8: Newton's Method, in which we found approximate solutions to equations like $x^2 - 2 = 0$ by repeatedly taking a linear approximation to $f(x) = x^2 - 2$. However, Newton's Method does not help to compute values of $\sin(x)$.

Common Taylor series

•
$$\frac{1}{1-x} = \sum_{n=0}^{\infty} x^n$$
 for $|x| < 1$ (Geometric Series).

•
$$\ln(1+x) = \sum_{n=1}^{\infty} (-1)^{n-1} \frac{x^n}{n}$$
 for $|x| < 1$.

•
$$(1+x)^p = \sum_{n=0}^{\infty} \frac{p(p-1)\cdots(p-n+1)}{n!} x^n$$
 for $|x| < 1$ (Binomial Series).

•
$$\exp(x) = \sum_{n=0}^{\infty} \frac{x^n}{n!}$$
 for all x .

•
$$\sin(x) = \sum_{n=0}^{\infty} (-1)^n \frac{x^{2n+1}}{(2n+1)!}$$
 for all x .

•
$$\cos(x) = \sum_{n=0}^{\infty} (-1)^n \frac{x^{2n}}{(2n)!}$$
 for all x .

Extra Topic: Irrationality of e. In §11.2, we saw that repeating decimals represent rational numbers (fractions), and every fraction can be written by long division as a repeating decimal. Thus, the non-repeating infinite decimals are the real numbers which cannot be written as fractions: they are *irrational*. However, it is difficult to prove that any given number (such as π or $\sqrt{2}$) is irrational.

We can use our series to prove the irrationality of the constant $e = 2.7182818284590\cdots$ To prove the negative proposition that e is not equal to any possible fraction $\frac{a}{b}$, we use the *method of contradiction*: that is, we assume that there *were* some fraction with $e = \frac{a}{b}$, and use this to deduce an impossible statement, which will show that the original assumption $e = \frac{a}{b}$ is also impossible.

Thus, using the Taylor series definition for e, we assume the possibility:

$$1 + \frac{1}{1!} + \frac{1}{2!} + \frac{1}{3!} + \cdots \stackrel{\text{def}}{=} e = \frac{a}{b}.$$

It is easy to show 2 < e < 3, so e is not a whole number, and would have denominator b > 1. Consider the bth order Taylor approximation:

$$1 + \frac{1}{1!} + \frac{1}{2!} + \frac{1}{3!} + \dots + \frac{1}{b!} + R_b = e, \qquad R_b = \sum_{n=b+1}^{\infty} \frac{1}{n!}.$$

We multiply by b! to clear denominators up to the $\frac{1}{b!}$ term:

$$b! + \frac{b!}{1!} + \frac{b!}{2!} + \dots + \frac{b!}{b!} + b!R_b = b!e = b!\frac{a}{b} = (b-1)!a.$$

The terms b!, $\frac{b!}{1!}$, $\frac{b!}{2!}$, ..., $\frac{b!}{b!}$ on the left are whole numbers, and (b-1)! a on the right is a whole number, so the remainder $b!R_b$ must also be a whole number. But it must also be very small, as we can see from a simple geometric series comparison:

$$b!R_b = \sum_{n=b+1}^{\infty} \frac{b!}{n!} = \frac{1}{b+1} + \frac{1}{(b+1)(b+2)} + \frac{1}{(b+1)(b+2)(b+3)} + \cdots$$

$$< \frac{1}{b+1} + \frac{1}{(b+1)^2} + \frac{1}{(b+1)^3} + \cdots$$

$$= \frac{1}{b+1} \frac{1}{1 - \frac{1}{b+1}} = \frac{1}{b+1} \frac{b+1}{b+1-1} = \frac{1}{b} < 1.$$

Thus, the same positive number $b!R_b$ would be both a whole number and less than 1, which is impossible. Thus the original assumption $e = \frac{a}{b}$ is also impossible.

[‡]We do not need the powerful Lagrange remainder formula here.